Design of crystal-detector assemblies for ChemiX spectrometer aboard Interhelioprobe

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Overview

- General informations about ChemiX
- Design of crystal-detector assemblies
- Results
- Conclusions

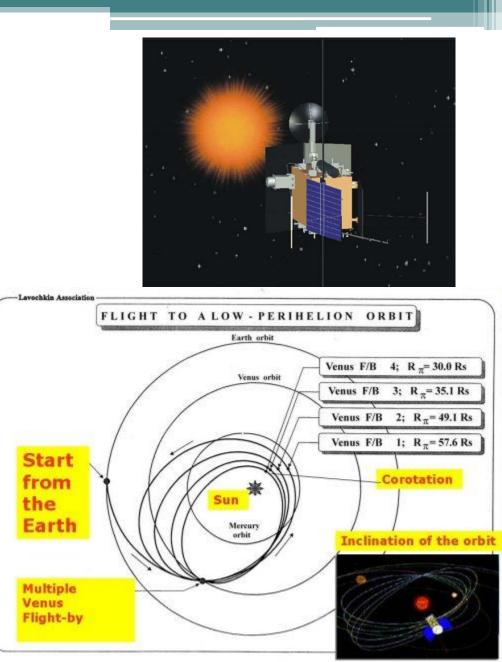
Interhelioprobe

Investigation of an inner Heliosphere and the Sun from close distances and from out-ofecliptic positions

Closest distance ~ 0.25 AU

Spatial resolution ~4 times better than from 1 AU

fluxes of the solar radiation ~15-20 times larger than in Earth vicinity.



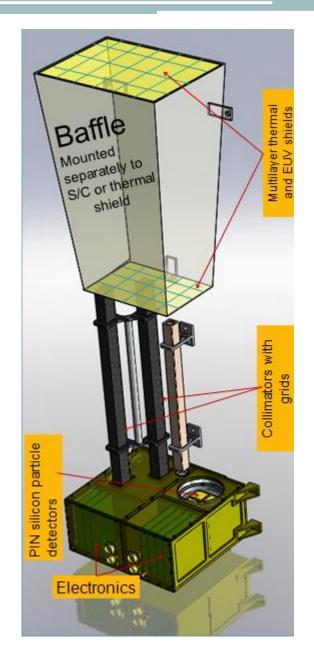
http://www.izmiran.ru

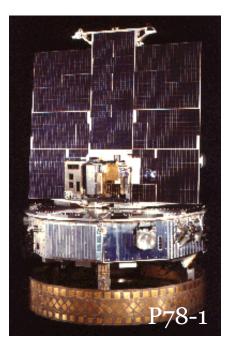
ChemiX

ChemiX is the Bragg crystal spectrometer for studies of Chemical composition of solar coronal plasmas based on measurements of X-ray spectra.

 $2 d \sin \theta = n \lambda$

The use of bent crystals allows the wider wavelength range to be integrated simultaneously, whereas a flat crystal spectrometer must scan to cover its wavelength range





Prewious missions









Prewious missions

launched on 31 July 2001

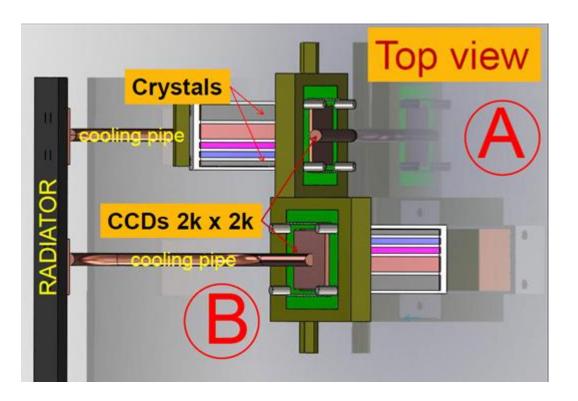


RESIK Bent crystal spectrometer

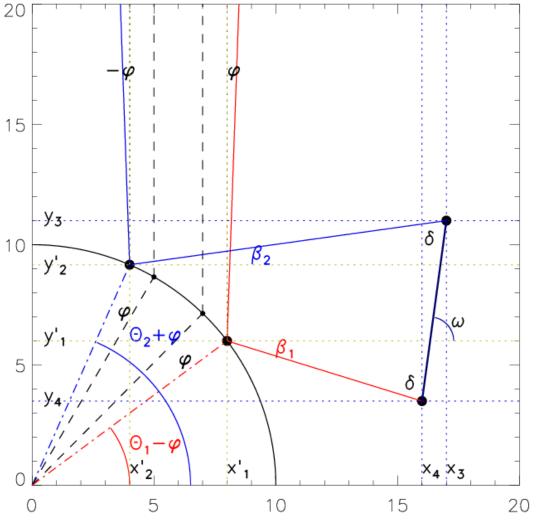


Diogeness Flat crystal spectrometer

Crystal-detector assembles



- two crystal-detector assemblies
- four wider mono-crystal wafers to cover the spectral range from approx. 1.5 A to 8.8 Å
- three dedicated crystal strips for the Dopplerometer in each crystal-detector assembly



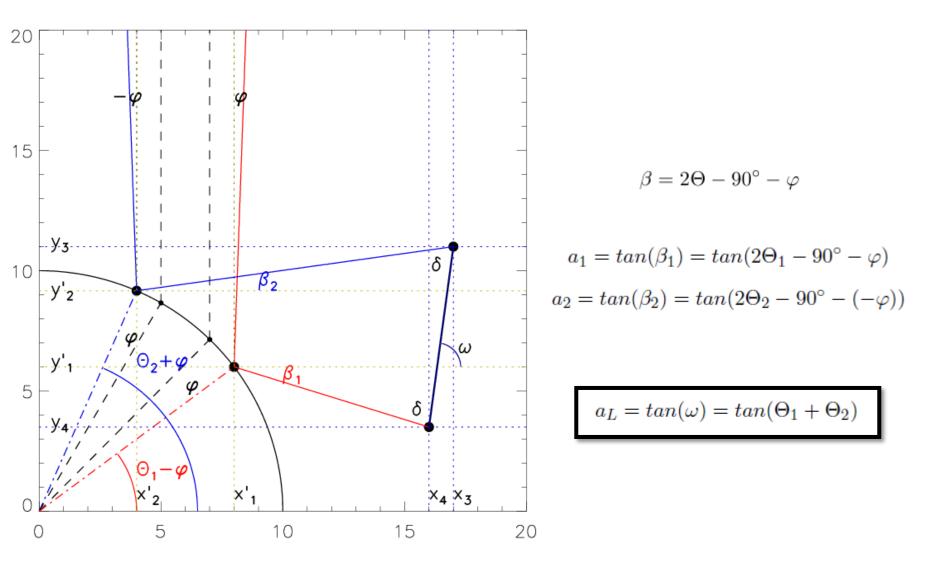
- (λ_1, λ_2) wavelenght range we want to observe
- (Θ_1, Θ_2) Bragg angles corresponding to the above wavelength
- φ offset, due to the fact that the Sun is not a point object
- R the radius of curvature of the crystal

$$x_1' = Rcos(\Theta_1 - \varphi)$$

$$y_1' = Rsin(\Theta_1 - \varphi)$$

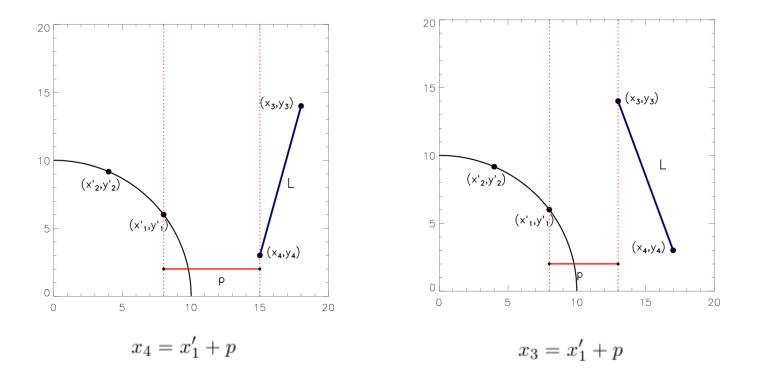
$$x_2' = R\cos(\Theta_2 - (-\varphi))$$

 $y_2' = Rsin(\Theta_2 - (-\varphi))$



L - detector lenght

p - minimum distance from the detector to the crystal



$$y_{3} = a_{2}x_{3} + b_{2}$$

$$y_{3} = a_{L}x_{3} + b_{L}$$

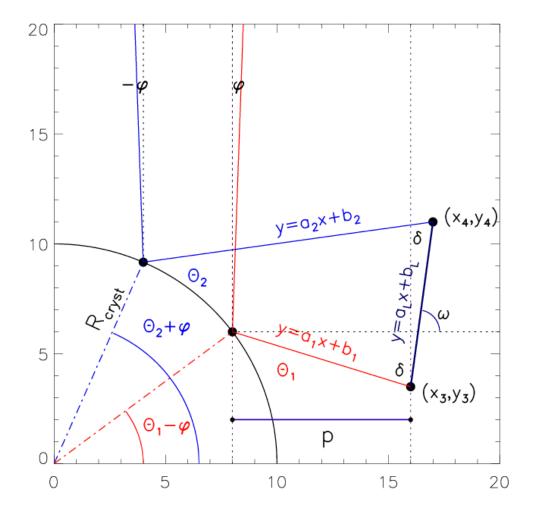
$$y_{4} = a_{1}x_{4} + b_{1}$$

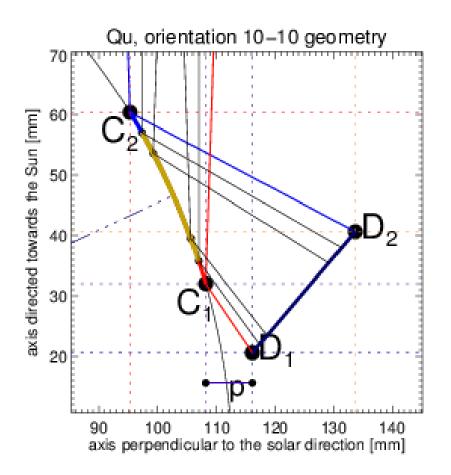
$$y_{4} = a_{L}x_{4} + b_{L}$$

$$L^{2} = (x_{4} - x_{3})^{2} + (y_{4} - x_{3})^{2}$$

$$a_{L} = tan(\omega) = tan(\Theta_{1} + \Theta_{2})$$

$$\begin{cases} x_{4} = x'_{1} + p \\ x_{3} = x'_{1} + p \end{cases}$$

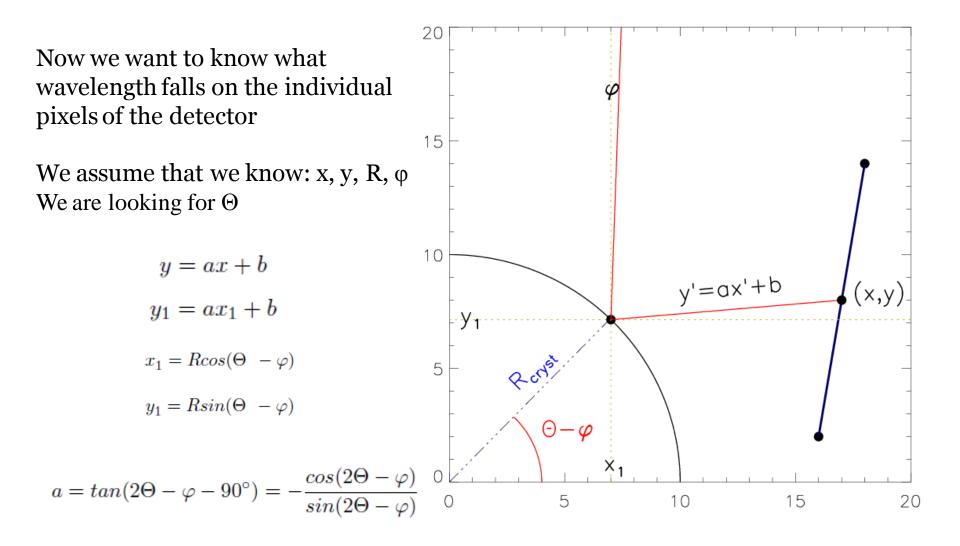




Waverange: 2.700 - 4.304 Å 2d: 8.514 Å

Radius of curvature: 112.789 mm

Detector slope: 48.88



$$y - y_1 = a(x - x_1)$$
$$y - Rsin(\Theta - \varphi) = -\frac{cos(2\Theta - \varphi)}{sin(2\Theta - \varphi)} [x - Rcos(\Theta - \varphi)]$$

 $cos^4 \Theta[4(A^2+B^2)] + cos^3 \Theta[-4RB] + cos^2 \Theta[-4(A^2+B^2)+R^2] + cos \Theta[2RB] + [B^2] = 0$

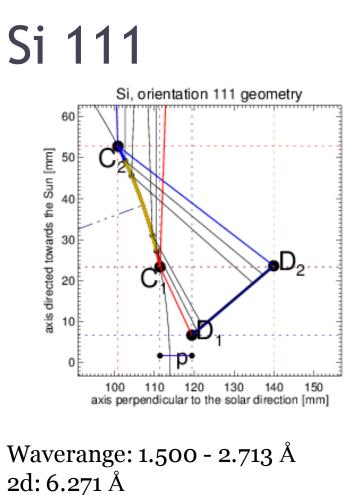
$$A = y cos\varphi + x sin\varphi$$
$$B = x cos\varphi - y sin\varphi$$

We get the equation of the fourth degree that we have to solve.

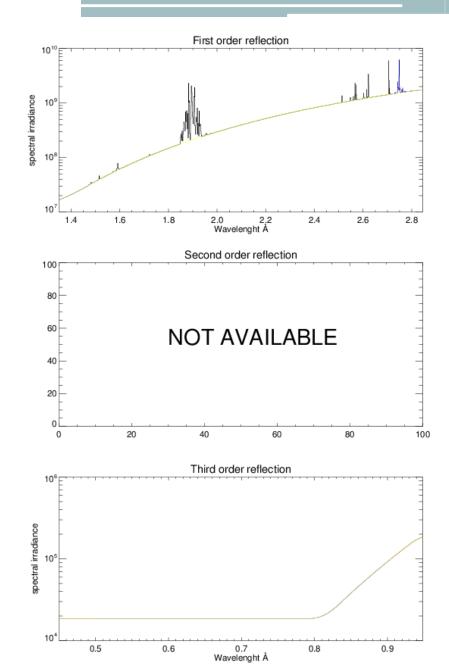
We can do this by using FZ_ROOTS function in IDL.

Selecting crystals

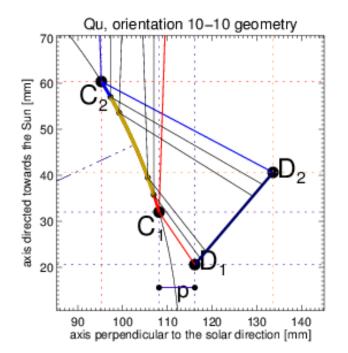
- we divide the observed wavelength range into four parts so that the bands overlap slightly
- we want to have some strong lines on those connecting areas
- we want to have most interesting lines on the middle of bands
- we don't want to observe fluorescent lines from crystal
- the radius of curvature of crystal can't be too small
- in the case of dopplerometr we want to have the most interesting lines observed with the best resolution



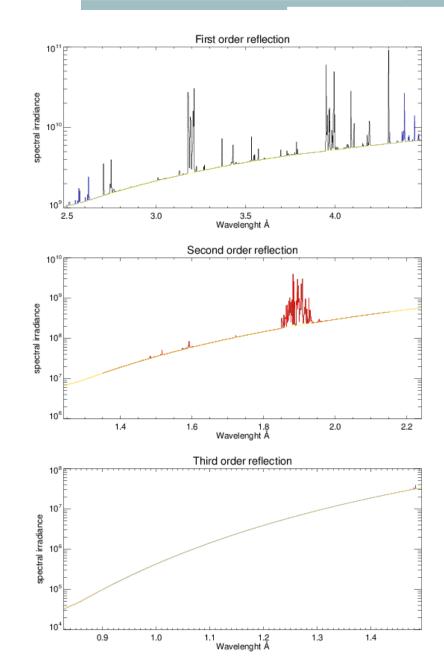
Radius of curvature: 113.84 mm Detector slope: 39.47



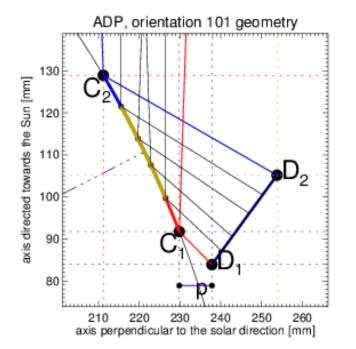
Quartz 10-10



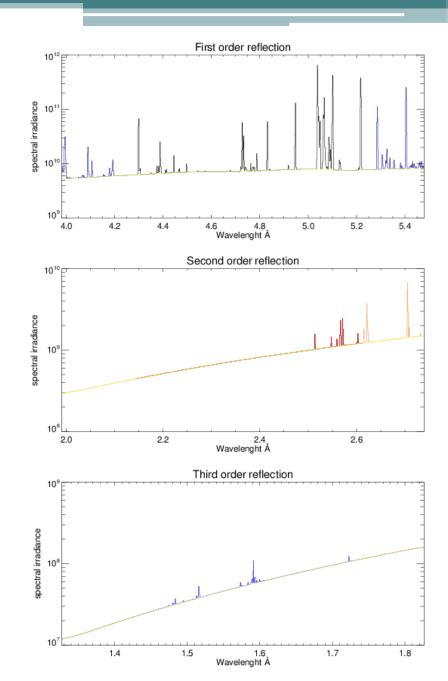
Waverange: 2.700 - 4.304 Å 2d: 8.514 Å Radius of curvature: 112.789 mm Detector slope: 48.88



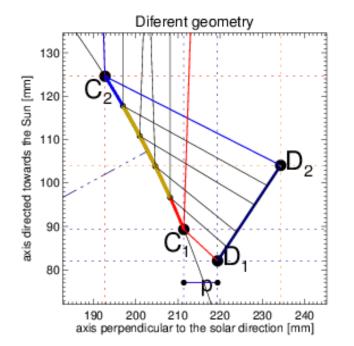
ADP 101



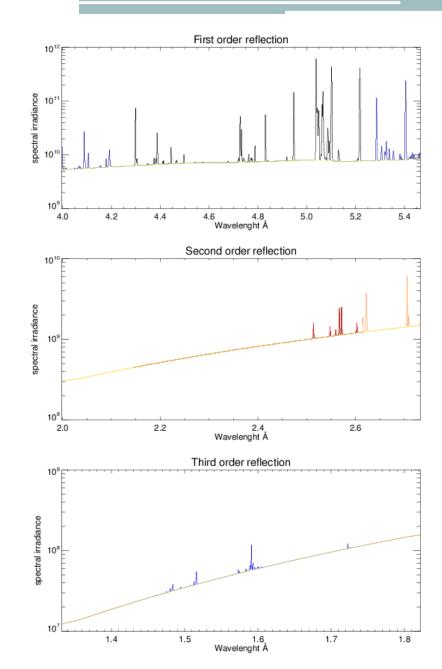
Waverange: 4.290 - 5.228 Å 2d: 10.648 Å Radius of curvature: 247.45 mm Detector slope: 53.16



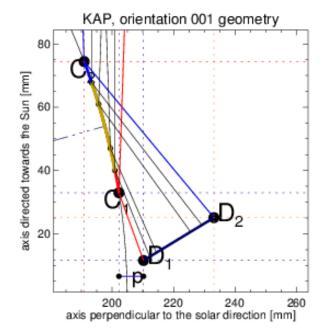
KDP 011



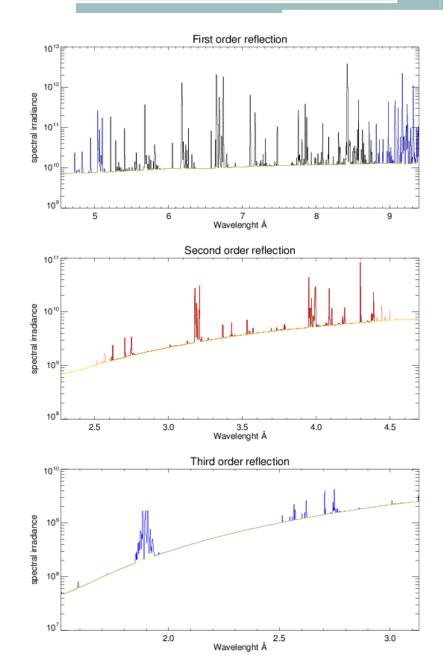
Waverange: 4.290 - 5.228 Å 2d: 10.185 Å Radius of curvature: 229.484 mm Detector slope: 56.52

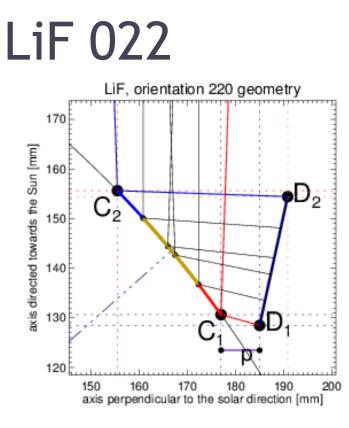




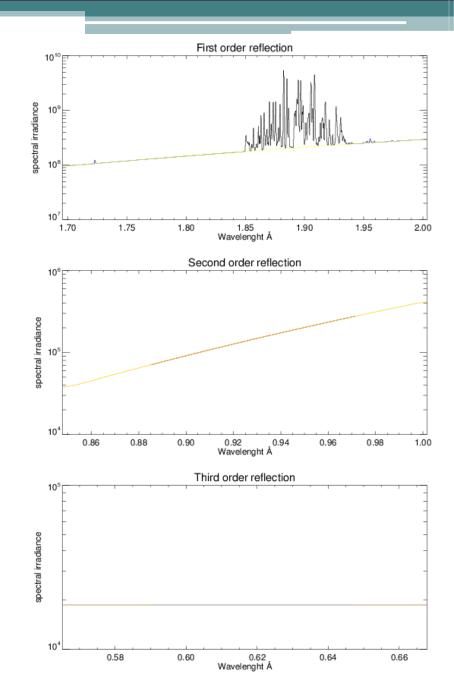


Waverange: 5.200 - 8.800 Å 2d: 26.64 Å Radius of curvature: 204.909 mm Detector slope:

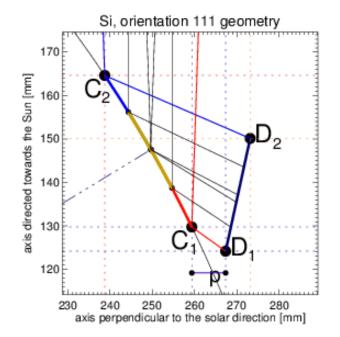




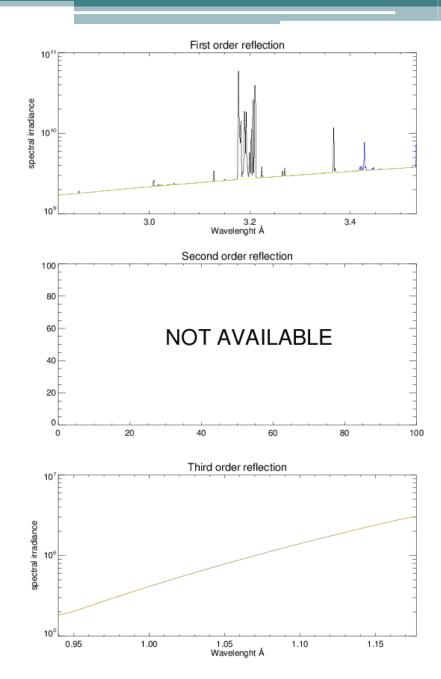
Waverange: 1.770 - 1.943 Å 2d: 2.848 Å Radius of curvature: 220.0 mm Detector slope: 77.35 (1.35 rad)



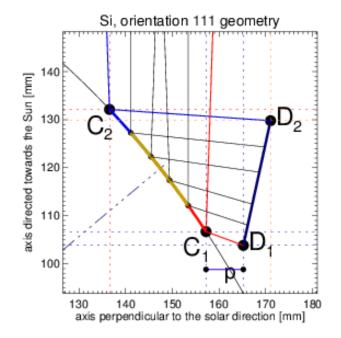
Si 111



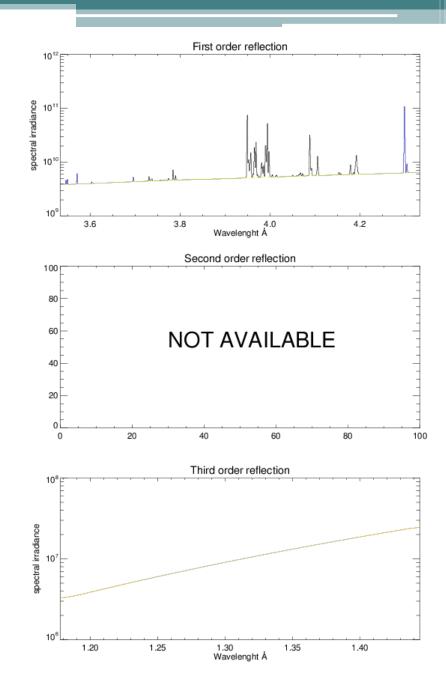
Waverange: 3.000 - 3.378 Å 2d: 6.271 Å Radius of curvature: 290.0 mm Detector slope: 77.35 (1.35 rad)



Si 111



Waverange: 3.700 - 4.201 Å 2d: 6.271 Å Radius of curvature: 190 mm Detector slope: 77.35 (1.35 rad)



No	crystal	orientati on	2d	wavelenght range [Å]	curvature radius [mm]	total desired crystal lenght [mm]	working crystal lenght [mm]	crystal width [mm]	number of crystals
Spectrometer									
1	Si	111	6.271	1.500 - 2.713	113.844	41.3	31.29	10	1
2	Quartz	10-10	8.514	2.700 - 4.304	112.789	41.2	31.17	10	1
3	ADP	101	10.648	4.290 - 5.228	247.454	51.6	41.61	10	1
	KDP	011	10.185	4.290 - 5.228	229.484	49.9	39.90	10	1
4	KAP	001	26.64	5.200 - 8.800	204.909	52.95	42.95	10	1
Dopplerometer									
1	LiF	022	2.848	1.770 - 1.943	220.000	43.0	32.98	10	2
2	Si	111	6.271	3.000 - 3.378	290.000	50.5	40.51	10	2
3	Si	111	6.271	3.700 - 4.201	190.000	42.8	32.78	10	2

Conclusions

- ChemiX will constitute the most advencet solar X-ray Bragg spectrometer ever flown
- It will contained 10 crystals bent to desired radius of curvature
- It will allow to observe the spectral range from 1.5Å to 8.8Å with very good temporal and spatial resolution
- X-ray Doppleromerer will allow to study of line-of-sight plasma motion in absolute reference system